

Stormier mid-Holocene southwest Indian Ocean due to poleward trending tropical cyclones

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Abstract

12 Geological evidence of past storminess is fundamental in contextualising long-term
13 climate variability and investigating future climate. Unlike the Atlantic and Pacific
14 basins, robust storminess reconstructions do not exist for most of the Indian Ocean,
15 despite the hazard tropical cyclones pose to the SE African margin. Here we combine
16 seismic stratigraphy with analysis of marine sediment cores to look for regionally
17 representative storm-related sediment deposits –or tempestites- intercalated in shoreface
18 sediments from the SW Indian Ocean off South Africa. Tempestites, represented by
19 hummocky seismic units, whose sediments have clear marine geochemical signatures,
20 are found to have been deposited between 6.5 and 4.6 cal kyr BP, when sea level was

21 between 0 and + 3 m above present. Deposition and preservation of the tempestites
22 reflect unprecedented tropical cyclone impacts, associated with periods of strongly
23 positive Indian Ocean Dipole (IOD) anomalies and linked to warmer sea surface
24 temperatures. Future climate projections suggest stronger positive IOD anomalies and
25 further intensification and poleward migration of tropical cyclones, like their mid-
26 Holocene predecessors. Given the rarity of tropical cyclone landfalls in South Africa,
27 this urges revaluation of hazards in areas along the southeast African coast likely to
28 become more vulnerable to landfalling tropical cyclones in future.

29 Palaeoclimatic reconstructions are vital in understanding past and future climate trends.
30 Because of the high impact of storms in coastal areas, climate projections often include
31 simulations of future storminess. Whereas many lines of evidence provide records of
32 past temperatures, pre-instrumental evidence of storminess is less abundant¹. Several
33 high-magnitude tropical storms (Hurricane Katrina, Cyclone Nargis, Hurricane Sandy,
34 Typhon Haiyan) in recent decades reveal the inadequacy of the instrumental record to
35 characterise storm recurrence intervals for exceptionally high-impact events. Evidence
36 of past storms and stormy periods is preserved in various marine geological proxies
37 including: (i) the paralic zone (washover deposits in back-barrier marsh sediments² and
38 erosional scarps in emergent barriers³), (ii) the shoreface and shelf (storm deposits or
39 tempestites)^{4,5} and (iii) deep ocean sediments (coarse-grained layers in pelagic
40 sequences)⁶. These storminess proxies have the potential to extend the instrumental
41 record if adequate chronological control can be established. To date, paralic and deep
42 ocean sediments have received most attention in this regard whereas shelf and shoreface
43 tempestites have long been recognised, but little used in palaeo-tempestology. Storm
44 deposits preserved in the shoreface or inner-shelf stratigraphy are effective tools to

45 assess the largest magnitude storms, however, their preservation potential is low as they
46 can potentially be reworked by subsequent storm events. When preserved, tempestites
47 provide a standing record of the largest storms, particularly intense tropical cyclones⁷,
48 and set a benchmark against which contemporary and future storminess can be assessed.

49 While the links between climate variability and tropical cyclone frequency, intensity
50 and track are nowadays better constrained⁸, uncertainties remain due to the limited
51 availability and quality of historical records and variations between modelling studies⁹.
52 Considering the sea surface temperature (SST) threshold (26.5°C) required for tropical
53 storms to develop¹⁰, ocean warming under a changing climate will lead to an expansion
54 of areas of tropical cyclone formation, consistent with the poleward displacement of
55 intensity maxima of tropical cyclones over the past decades¹¹. However, there is low
56 confidence in projected changes for tropical cyclone genesis, track and duration, despite
57 the likely decrease in frequency and increase in intensity⁹, and there is particularly low
58 confidence in basin-specific projections of storminess and associated storm surges¹².

59 Part of the reason for this uncertainty is the lack of palaeo-tempest records against
60 which to compare climate model outputs, which need to be extended both in time and
61 space⁸. Various records of palaeo-tempests from the Pacific and Atlantic Oceans^{7,13}
62 have been linked to modes of climate variability such as the NAO, ENSO or the PDO¹⁴.
63 No such records of palaeo-tempests have been reported from the coasts of the SW
64 Indian Ocean despite the known impacts of regular tropical cyclones^{15,16,17,18}. Here we
65 present evidence from tempestites preserved on the lower shoreface off Durban, South
66 Africa (29.9° S, 31.0° E) (Fig. 1), which record a period of enhanced storminess during
67 the mid-Holocene. We assesses the timing, genesis and preservation of the tempestites

68 and their association with tropical cyclones and climate variability for the SE African
69 coast, providing a benchmark for future assessment and modelling of tropical cyclone-
70 climate links¹⁹ in the Indian Ocean.

71

72 Shoreface tempestites in the SW Indian Ocean

73 Along the microtidal, wave-dominated east coast of southern Africa, which is exposed
74 to a range of tropical and extratropical cyclones driving extreme storm waves, sea level
75 has risen episodically over the last 18 kyrs^{20,21}. It reached the present level ~ 6 kyrs BP,
76 after which two minor highstands (+3.5 and +1.5 m) occurred at 4.5 and 1.6 kyrs BP
77 (Fig. 1b). Offshore Durban (South Africa), the lower shoreface (between fair- and
78 storm-weather wave base²²), is characterised by Holocene unconsolidated transgressive
79 sediments that mantle and abut occasional aeolianite pinnacles²³. Here, we present the
80 seismic stratigraphy and age-controlled sedimentary and geochemical analysis of two
81 cores (see Methods) collected during RV METEOR Cruise M102²⁴.

82 The sedimentary succession intersected by cores GeoB18304-1 and GeoB18303-2
83 (Fig.1) comprises three units (1-3) that overlie and postdate the Holocene wave
84 ravinement²³ (Fig. 2). The succession imaged occupies the mid-shelf where the
85 sediment cover is thin and patchy, each unit separated in space but necessarily in time.
86 The contemporary depth of storm wave sediment reworking is estimated at ~ 40 m²⁵,
87 just below which these three units occur. The sandy nature and position of the units
88 identify them as part of the contemporary lower shoreface²³.

89 The lower shoreface comprises a seaward Unit 1 that consists of irregular, wavy to
90 chaotic high-amplitude reflectors (e.g. Fig. 2a, Extended Data Fig. 1a). Unit 2
91 comprises two facies: a proximal set of flat-lying reflectors that become progradational
92 with depth (2A) (Extended data Fig. 1a), overlapped by hummocky, wavy to irregular and
93 chaotic reflectors (2B, Fig. 2b), all of which are truncated by Surface ii (Extended data
94 Fig. 1b, Fig. 2). This surface is irregular and is overlain by Unit 3, which comprises
95 another series of irregular, wavy to chaotic high-amplitude reflectors and forms the
96 shallowest accumulation of the lower shoreface (Fig. 2b). Units 1 and 2 are separated by
97 zones of non-deposition, marked by exposed erosional surfaces and aeolianite pinnacles
98 at -55 to -60 m (Extended data Fig. 1).

99 The seaward core (GeoB18303-2) in 60 m water depth contains a uniform succession of
100 shelly, medium to coarse sands (Fig. 3a). Seismic unit 1 is intersected by the upper 3.5
101 m of the core. This unit is characterised by the presence of several mudballs between
102 1.8 and 3 m depth. These date from 12 052 cal yr BP in the lower sections followed by a
103 significant hiatus between 11 224 cal yr and 4177 cal yr BP when the most recent
104 deposition of mudballs occurred (Fig. 3a).

105 The landward core (GeoB18304-1) in 35 m water depth shows a general fining-upward
106 succession from a series of pebbly coarse sands to medium sand that correlates to Unit
107 2B (Fig. 3b). The lower portion of the core comprises a series of coarse grained, sharp-
108 topped and sharp-based event beds. No datable material was found at their upper
109 boundary, but ages below and above date from 6980 cal yr BP to a minimum of 2619
110 cal yr BP (Fig. 3b). The uppermost part of the core correlates with Unit 3 and comprises
111 a coarsening upward succession of coarse to very coarse sands.

112 The most significant changes in grain size and element concentration in core
113 GeoB18303-2 occur between 1.5 m and 2.25 m (Fig. 3c and d). Here significant
114 decreases in the concentration of the marine fraction elements including Ca (103.76
115 g/kg) and Sr (174 mg/kg) are evident, with corresponding finer grain sizes (Fig 3c).
116 Associated with these depths are increases in the terrigenous fraction elements including
117 Si (204.91 g/kg), Al (41.57 g/kg), K (15.97 g/kg), Ti (319 mg/kg) and Rb (7 mg/kg), as
118 well as an increase in Fe (20.15 g/kg). These coincide with the matrix that hosts the
119 mudballs (Fig. 3d).

120 Grain size and elemental concentrations in core GeoB18304-1 vary little with depth
121 until 3.15 m (Fig. 3e and f). From 3.15 m down to the basal layers, there is a significant
122 scatter with multiple switching between high and low concentrations. There are multiple
123 spikes in abundance of the marine elements towards the core base (Fig. 3f). The
124 increases in marine elemental abundances are associated with the coarsest grain sizes
125 that form the base of the small-scale, fining-upwards packages (Fig. 3b).

126 Potential mobilization of seafloor sediments based on modelled bed shear stress during
127 extreme storm waves offshore Durban (see Methods) indicate that coarse sand, the most
128 common material found in both cores, is mobilized over the entire domain (Extended
129 Data Fig. 2a). The thresholds for mobilization of gravel-sized sediments (Extended Data
130 Fig. 2b), the coarsest material found in the proximal cores, are similarly exceeded along
131 the entire lower shoreface. For the 100 yr return-period storm, the entire shoreface and
132 inner shelf would be subject to disturbance for both classes of coarse sediment
133 (Extended Data Fig. 2c,d).

134

135 Storm deposits

136 The shoreface units (1-3) post-date the early Holocene wave ravinement surface
137 identified by previous authors²³. Unit 1 onlaps the various aeolianite/beachrock ridges
138 as a series of seaward-thinning wedges of shelly sediment with notably irregular, wavy
139 to chaotic high-amplitude reflectors. Distally, unit 1 comprises mudballs within a very
140 coarse sand matrix. The terrestrial origin of the mudballs is indicated by high Ti
141 abundance (Fig. 3d). They occur within a coarse-grained shell hash with high marine
142 elemental signatures. Mudballs on the shelf are commonly found in storm-dominated
143 settings where the coastline is undergoing transgressive erosion. They are derived
144 through storm-driven erosion of muddy coastal/fluvial sediments and subsequent
145 offshore transport in storm-return flows that extend below storm wave base^{26,27}. The
146 mud is likely derived from an outcropping or subcropping source on the adjacent
147 foreshore. This occurs presently in the study area, when storm erosion exposes laterally
148 continuous back-barrier mud layers along the shoreline²⁸. Based on their terrestrial
149 signatures and transgressively eroding setting, we consider the mudballs to represent
150 similar storm-based erosion of terrestrial-sourced muds from the foreshore and
151 subsequent deposition within the tempestite sequence in the lower shoreface, as a result
152 of storm return flows. No further mudballs occurred in the upper stratigraphy of either
153 of the cores. The dated outer layer of the mudball (4117 cal yr BP) reflects the
154 maximum age of deposition of this material on the shelf.

155 Unit 2 is present at depths from 60 to 40 m (Extended Data Fig. 1), with isolated
156 pockets of sub-Unit 2B occurring at the termini of the prograding sub-Unit 2A. The
157 high abundance of marine fraction elements, separated by finer material with high

158 terrigenous elemental abundance are indicative of periodic high-energy marine
159 events^{29,30}. The marine-dominated shell and pebble hash horizons are similar to deposits
160 (“rippled scour depressions”) associated with storm scour on the inner shelf^{31,32,33}. The
161 small-scale, sharp-based coarse packages that terminate with terrestrial element-rich
162 sands are similar to the tempestites described by others²⁶. Sub-Unit 2B is thus
163 considered to comprise a series of storm-generated gravel/sand couplets.

164 Dates from the overlying Unit 3 constrain the deposition of the overlying shoreface
165 sediments to 2619 cal yr BP and 1878 cal yr BP. Units 1 and 2B, and the storm intervals
166 they record, span two distinct time periods. The distal storm deposits (mudballs in storm
167 return flow deposits) date from 12 052 cal yr BP to 11 224 cal yr BP, followed by a
168 hiatus, to 4 177 cal yrs BP. The more proximal storm deposits more closely match this
169 younger date, and span the 6 980 cal yr BP to 4910 cal yr BP interval. In the context of
170 palaeo-sea levels, the timing of deposition of the older tempestites is associated with a
171 time when sea levels were ~ 30-45 m below mean sea level^{21,34} (Fig. 1), whereas the
172 proximal group occurred when sea level was between 0 and + 3 m²¹ (Fig. 1). The older
173 and distal storm deposits were initially associated with a lowered wave base (~ -45 to -
174 60 m from 12 to 11 ka), at which time and based on their depths, they likely developed
175 in upper shoreface-hosted rippled scour depressions. As sea level rose to the present,
176 periodic storm deposition continued on the outer shoreface until 4 177 cal yr BP.

177 The proximal deposits relate entirely to deposition below storm wave base under
178 contemporary sea level conditions²⁵. Preservation potential of tempestites is low
179 because subsequent storms rework older deposits³⁵, but intense tropical cyclones
180 generate thick shoreface deposits that can survive physical and biological reworking³⁶.

181 While the largest of contemporary storms recorded in the coast of Durban appears
182 capable of remobilising gravel-sized particles over the entire lower shoreface (Extended
183 Data Fig. 2), the tempestite horizons are still preserved in the substrata. The storm
184 deposits thus appear to record events of a magnitude that has not been exceeded since.
185 We attribute this to intense tropical storms given the geographical position of Durban in
186 relation to the Southern Indian Ocean tropical cyclone belt³⁷. In the overlying
187 succession of shoreface sediments, there are no further storm event horizons, suggesting
188 no further impingement by intense storms capable of forming such pervasive
189 tempestites on the seabed.

190 Similar sequences of tempestites have been associated with centennial to millennial
191 periods of increased tropical cyclone activity in the Atlantic Ocean^{7,38}, produced by
192 landfall of hurricanes of category 3 and higher. While tropical cyclones of such intensity
193 have not made landfall along the eastern coast of South Africa in the past 5
194 decades^{18,37,39}, and less intense (category 1 and 2) tropical cyclones rarely make landfall
195 along this coastline⁴⁰, the tempestites archived in the cores and seismic stratigraphy
196 point to a prolonged mid-Holocene period of very intense tropical cyclone activity in
197 southern Africa.

198

199 Paleo-climatic context

200 The mid-Holocene tempestite record of core GeoB18304-1 indicates that intense storm
201 activity started at or before the oldest date the core (6980 cal yrs BP) and was ongoing
202 at least until the youngest date (4816 cal yrs BP). Studies elsewhere suggest that intense
203 storminess is likely to be associated with increased regional SST⁴¹, which in the western

204 Indian Ocean is related to the IOD⁴². The IOD is considered a major climatic driver
205 across the Indian Ocean region throughout the Holocene⁴³. Positive IOD events are
206 associated with greater-than-average SST in the western Indian Ocean and increased
207 rainfall over East Africa⁴⁴. Positive IOD anomalies occur when strong easterly winds
208 and weakening of eastward oceanic currents along the equatorial Indian Ocean facilitate
209 atmospheric and oceanic current reversals^{45,46,47}. The majority of studies of atmospheric
210 and oceanic circulation in the Indian Ocean link rapid SST warming in the west to
211 strong easterly winds and weakening of eastward oceanic currents along the equatorial
212 Indian Ocean. Enhanced convection over the Indian Ocean reflects a positive IOD
213 anomaly⁴⁸. Large changes in the monsoon rainfall in the eastern Indian Ocean have been
214 attributed to the occurrence of strong positive IOD anomalies^{49,50}, during which SST is
215 high and the likelihood of intense and more frequent tropical cyclones in the western
216 Indian Ocean increases. Strong positive IOD induces extreme weather events in eastern
217 Africa⁵¹, and is associated with increased rainfall along the coasts of Mozambique and
218 South Africa⁵².

219 This period coincides with strong positive IOD events that caused aridity and SST
220 cooling over the eastern Indian Ocean, while the western margin experienced increased
221 precipitation and positive SST anomalies^{53,54,55,56}. When compared to Mauritian climate
222 records (Fig. 4b), the tempestite deposition matches an overall period of negative IOD
223 state with strong positive anomalies⁵⁷. This period is further correlated with records
224 offshore Somalia⁵⁸ and Tanzania⁴³ which reveal warmer SST for the western Indian
225 Ocean between 7.8 and 4.7 ka BP (Fig. 4c). Higher SST not only increases the
226 likelihood of intense and more frequent tropical cyclones, but also contributes to a
227 southward shift in the latitudinal position of the 26 °C and 27 °C isotherms, and

228 potential changes in the location of tropical cyclone landfalls, tracking south of
229 Madagascar and making landfall in higher latitude regions along the coasts of
230 Mozambique and South Africa¹⁶. After 4.3 ka the lack of tempestites is also associated
231 with a shift towards a stronger El Niño and a less prominent Eastern Indian Ocean
232 monsoon since 3600 BP⁵⁴ (Fig. 4d).

233 Examinations regarding changes to tropical cyclone frequency and intensity over the
234 southern Indian Ocean under a warming climate have been inconclusive and often
235 contradictory^{9,17,39,59,60}. However, an increasing trend in the intensity and duration of
236 tropical cyclones associated with warming SST and upper ocean heat content in the
237 southern Indian Ocean has been observed in the last two decades¹⁸. Under high
238 greenhouse emission scenarios, multi-model climate projections robustly indicate more
239 frequent⁶¹ and more intense⁵¹ strong positive IOD events, driven by increased SST
240 variability in the western Indian Ocean. Therefore, global warming will likely lead to
241 enhanced storminess in Southern Africa, linked to strong positive IOD events associated
242 with more intense and southward tracking tropical cyclones, of which the mid-Holocene
243 deposits on the Durban shelf provide a clear analogue.

244

245 These findings demonstrate the potential of shoreface deposits as a proxy for past
246 storminess and intense tropical-cyclone landfall. Two phases of enhanced storminess
247 are recorded. One is associated with an early Holocene sea-level of ca. -40 m and is
248 preserved in a drowned shoreface. The second (6.9 to 4.8 ka) is associated with
249 contemporary sea-levels and records a period of enhanced storminess that, alongside
250 other proxies, evidences a clear association with strong positive IOD events. Higher

251 SST and strong positive IOD events due to global warming are likely to lead to more
252 intense, frequent and southward tracking tropical cyclones, whose impacts will be
253 significantly greater than those of the present and the historic past along the coast of
254 southern Africa.

255

256 Author Contributions Statement

257 AG led the paper conceptualisation, data collection, analysis, figure drafting, and
258 together with JAGC managed the paper writing and editorial review. SD performed the
259 laboratory analyses and figure drafting. CL performed the modelling and assisted in
260 data analysis, writing, figure drafting and editorial review. AH and MZ assisted with
261 data collection, writing, editorial review, with MZ the principal funding recipient.

262 Competing Interests Statement

263 The authors declare no competing interests.

264 Data availability

265 Samples and data (inorganic data, radiocarbon analyses) are respectively archived at the
266 GeoB Core Repository and Pangaea (www.pangaea.de) both located at MARUM,
267 University of Bremen. Modelling results are available on request of the corresponding
268 author.

269 Figure captions

270 **Figure 1.** Location of the Durban shelf and study site with multibeam bathymetry²⁰
271 (courtesy eThekweni Municipality), seismic coverage (grey lines) and core sites. Inset
272 **b**, SE African sea level curve²¹. SA=South Africa, Moz = Mozambique, Tan =

273 Tanzania, Ken = Kenya, Som = Somalia, Sey = Seychelles, Maur = Mauritius. Map
274 projection WGS84, UTM 36S

275 **Figure 2.** Zoomed in ultra-high-resolution seismic stratigraphy of the lower shoreface.
276 **a)** core site GeoB18303-2. **b,** GeoB18304-1. Note the hummocky nature of unit 2B
277 intersected by GeoB18304-1. Profile positions of a and b denoted in figure 1 and the
278 full profiles are provided in Extended Data Fig. 1. WRS = wave ravinement surface

279 **Figure 3.** Downcore variations and chronology. **a,** GeoB1803-2. **b,** GeoB18304-1.
280 Areas of interest are outlined by shaded grey boxes. WRS = Holocene wave ravinement
281 surface. **c,** bulk sediment grain size variations GeoB1803-2. **d,** downcore elemental
282 distributions GeoB1803-2. **e,** bulk sediment grain size variations GeoB1804-1. **f,**
283 downcore elemental distributions GeoB1804-1. Grey lines link spikes in grain size with
284 corresponding peaks and troughs in marine and terrestrial material. Cl = clay, Si = silt,
285 VFS = very fine sand, FS = fine sand, MS = medium sand, CS = coarse sand, VCS =
286 very coarse sand, P = pebbles, Gr = granules, Co = cobbles, B = boulders

287 **Figure 4.** Lithologic and geochemical variations compared to major climatic
288 oscillations in the South West Indian Ocean (SWIO). **a,** downcore variations of grain
289 size, Ca and Ti abundances and geochronology of GeoB18304-1, **b,** fluctuating Ca/Ti
290 ratios in cores from Mauritius⁵⁷, **c,** SST anomalies (lines) from Tanzania⁴³ and
291 reconstructed alkenone palaeothermometry SST data from Tanzania (circles)⁵⁸, **d,** El
292 Niño events per 100 years⁶². Red circles are strong El Niño Indian Ocean Dipole (IOD)
293 events, blue circles are strong monsoon IOD events, grey blocks denote period of
294 interest. ENSO = El Niño-Southern Oscillation

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503

504

505 **Extended Data Fig. 1.** Seismic reflection profiles and interpretations of the seismic
506 stratigraphy of the Durban shelf. **a**, full record including figure 2a. **b**, full record
507 including figure 2b.

508 **Extended Data Fig. 2.** Bed shear stress represented according to the thresholds for
509 sediment mobility. Model results for the largest recorded storm offshore Durban for: **a**,
510 coarse sand, **b**, fine gravel, and the 100 yr return-period storm for **c**, coarse sand, **d**, fine
511 gravel. Areas below threshold are blanked. Note that at the GeoB18304-1 site, granule-
512 size sediment would be mobilised, but not at the GeoB18303-2 site.

513 **Methods**

514 **Regional setting:** The eastern coast of South Africa is characterised by mean annual
515 significant wave height of 1.65 m⁶³, and spring and neap tidal ranges are between ~ 1.8
516 m and ~0.5 m, respectively⁶⁴. Extreme waves in this coastal area are driven by tropical
517 cyclones, mid-latitude (extratropical) cyclones and cut-off lows^{63,65}. If tropical cyclones
518 become stationary south east of Madagascar they can drive large wave events along the
519 east coast of South Africa^{63,65}, while cut-off low systems may also drive large waves
520 storm waves and surges. No tropical cyclones made landfall in the coast of South Africa
521 since wave records began in the early 1980's, but an intense cut-off low system
522 occurred in March 2007. The storm generated the largest waves recorded, with peak
523 significant wave height of 8.5 m, corresponding to a return-period of 32 to 61 years⁶⁶.
524 This event caused widespread coastal erosion and infrastructural damage²⁵.

525 **Geophysical surveying and coring:** The shallow sub-surface geology was examined
526 using ultra-high-resolution 0.5kHz PARASOUND collected during RV METEOR
527 Cruise M102 in December 2013²⁴. All data were processed by high and low band pass

528 filtering and gain application and exported as SEG-Y data for visualisation in the
529 Kingdom Suite software package. The processed PARASOUND data resolve to ~ 10
530 cm in the vertical domain with a maximum penetration of ~ 20 m in localised areas. In
531 all lines, the upper 5 m of the seafloor sediment package were resolved with a high level
532 of detail. Key targets were identified from the ultra-high-resolution seismic packages for
533 coring. Three vibrocores were collected during the same cruise²⁴, two of which
534 (GeoB18303-2 and GeoB18304-1) are described in this study (Fig. 1). Previous
535 descriptions of the seismic stratigraphy are included together with those of this study in
536 Supplementary Table 1. Multibeam bathymetry⁶⁷ collected by the eThekweni
537 Municipality were integrated with the ultra-high resolution seismic and core data in
538 order to assess the spatial distribution of tempestite signatures on the lower Durban
539 shoreface.

540 **Laboratory analysis:** Cores were split onboard and logged according to standard
541 sedimentological procedures. Sub-sampling at 5cm intervals for grain size and
542 geochemical analyses was undertaken, together with sampling for material suitable for
543 Accelerator Mass Spectrometry (AMS) ^{14}C dating. A total of 13 samples were
544 collected from cores GeoB18303-2 and GeoB18304-1 for dating purposes. The material
545 used for AMS ^{14}C dating is listed in Extended Data Fig. 3. All shell material was
546 selected from in-situ life position, especially in the case of bivalves that were still
547 articulated. Wherever possible, the most intact shells were chosen with the least amount
548 of bleaching of the shell exterior. All dates are corrected for reservoir effect with a ΔR
549 of 121 ± 16 ^{14}C yr⁶⁸. The dates discussed in this manuscript are median values; the two
550 sigma ranges are indicated in Supplementary Table 2.

551 Particle size analysis was undertaken for both the bulk and terrigenous sediment
552 fractions. The samples were sieved to obtain the bulk grain size distribution with the
553 result of the analysis presented as phi values where the mean, median, sorting and
554 skewness were calculated using the Folk and Ward equations. For the terrigenous grain
555 fractions, the sediment samples were treated with 10% HCl, H₂O₂ and NaOH to remove
556 calcium carbonate, organic matter, and biogenic opal, respectively. The samples were
557 then suspended in demineralised water with the addition of Na₄P₂O₇ to prevent the
558 formation of aggregates. The particle size distribution was measured with a Coulter
559 laser particle sizer LS 13 320 (MARUM, University of Bremen, Germany) generating
560 92 size classes from 0.4 to 2000 µm. For this study the mean grain-size data are
561 displayed as phi.

562 Additional samples were collected at selected locations corresponding to significant
563 results obtained from the grain size analyses. Sample pre-treatment consisted of drying
564 and grinding for 120 seconds in a silicon nitride vessel to prevent contamination
565 (Planetary Micro Mill PULVERISETTE 7 premium line, MARUM, University of
566 Bremen, Germany), to assure that all particles were smaller than 63 µm.

567 Elemental compositions were measured on 205 sediment samples where 4 grams of
568 each sample compressed at 25 kPa, were used to analyse for major, minor and trace
569 element composition by X-Ray Fluorescence spectrometry (Panalytical epsilon 3 XL,
570 Bremen University, Germany). USGS and Chinese rock and sediment standard
571 reference material GBW 07316 was measured simultaneously and gave results within +/-
572 3-5% of certified values.

573

574 **Wave modelling and sediment mobility analysis:** Shoreface sediment mobility in
575 response to storm wave forcing was analysed using the nearshore wave propagation
576 model SWAN version 41.20AB^{69,70}. Simulations of the wave field were performed for
577 the maximum wave conditions during the largest storm recorded offshore Durban⁶³
578 (March 2007; significant wave height of 8.5 m and peak wave period of 16.6 s) and for
579 the 100-year return period storm⁶⁶ (significant wave height of 10.3 m and peak period of
580 17.4 s). SWAN is a depth and phase-averaged, third-generation wave model that
581 simulates de refractive propagation and evolution of the wave spectrum. The model was
582 run in stationary mode, i.e. time is removed from the computations and waves are
583 assumed to propagate instantaneously across the modelling domain, using default
584 parameters in order to account for bottom friction dissipation, non-linear wave
585 interaction, diffraction and white-capping dissipation⁷¹. A regular structured grid with 5
586 meters resolution was used for representing the computational domain, matching the
587 bathymetric grid used to represent the bottom conditions (Fig. 1).

588 Considering the dependency of near-bed sediment movement on the bottom orbital
589 velocity amplitude⁷², outputs from SWAN included the root-mean-square of the orbital
590 motion near the bottom (U_{rms}) for the entire modelling domain, computed considering a
591 JONSWAP spectral shape and empirical bottom friction model and linear wave
592 theory⁷³. To evaluate the potential for wave-induced coarse sediment entrainment and
593 transport during modelled storm conditions, the threshold bed shear stress for initiation
594 of sediment transport (T_{cr}) based on the modified Shields parameter was computed⁷² for
595 coarse sand (d_{50} = 0.5 to 2mm) and fine gravel (d_{50} =2 to 8 mm)⁷⁴. T_{cr} values of 0.63
596 N/m² and 4.00 N/m² were obtained for the mean class values of coarse sand (d_{50} = 1.25
597 mm) and fine gravel (d_{50} = 5 mm), respectively.

598 These values were then compared to the spatially variable bed shear stress under waves
599 (T_{ws}), considering that on a flat, non-rippled bed typical of coarse sediments, the bed
600 shear stress can be simplified and only the wave-skin friction component (T_{ws}) is
601 required to determine the hydrodynamic forcing acting on the bed and driving sediment
602 entrainment and transport⁷². T_{ws} was computed using modelled bottom orbital velocity
603 ($U_w = U_{rms}$) and the wave friction factor (f_w) according to:

604
$$T_{ws} = \frac{1}{2} p f_w U_w^2$$

605 where p is seawater density (1027 kg/m³), U_w corresponds to U_{rms} modelled with
606 SWAN and f_w computed using the formulation⁷²:

607
$$f_w = 1.39(A/z_0)^{-0.52}$$

608 where A is the semi-orbital excursion ($U_w T / 2\pi$), and z_0 the bed roughness length
609 ($d_{50}/12$).

610

611 **Data availability**

612 Seismic and core data (geochemical, grain size and chronology) are available at
613 Pangaea (www.pangaea.de). Modelling data are available on request from AG or CL.

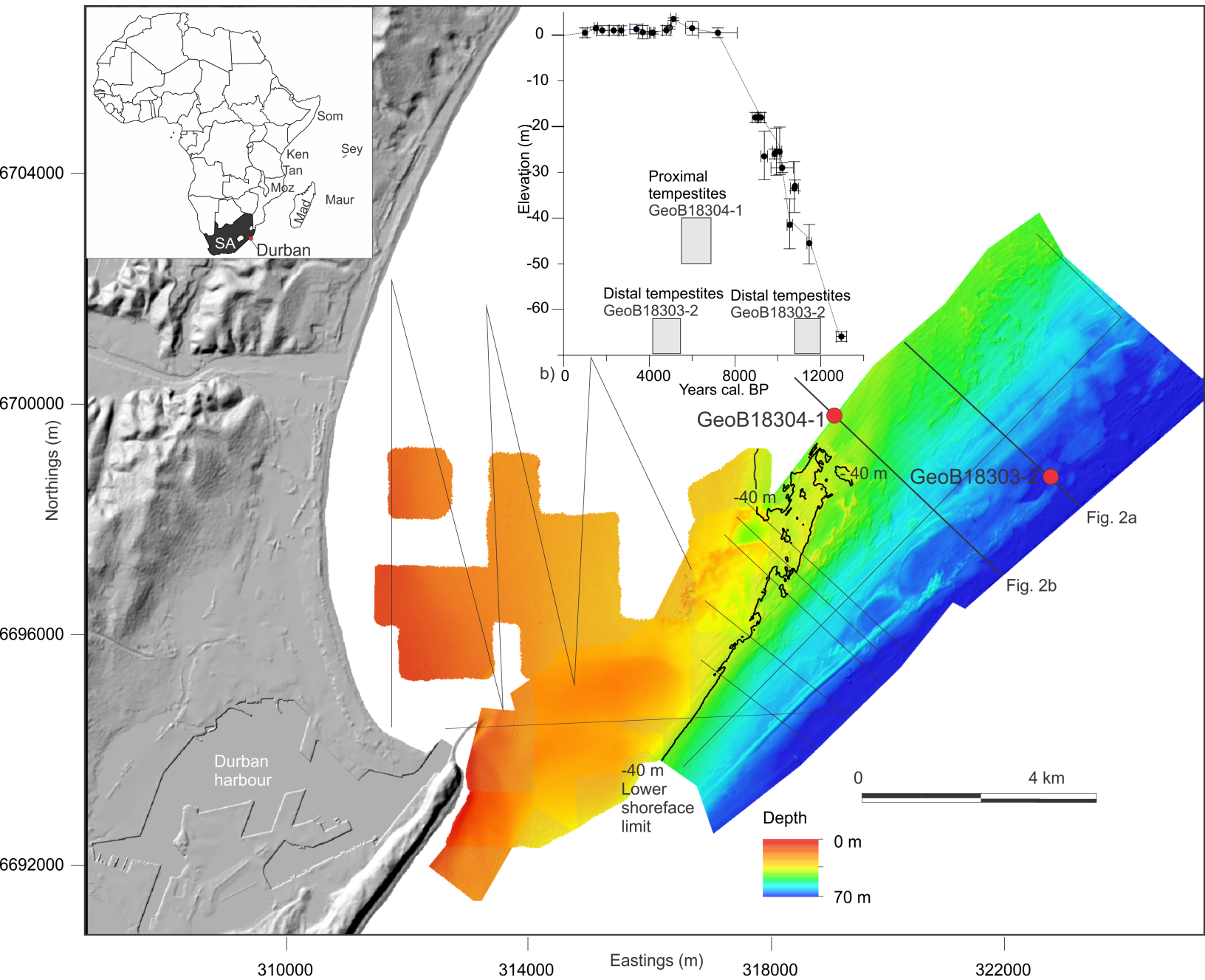
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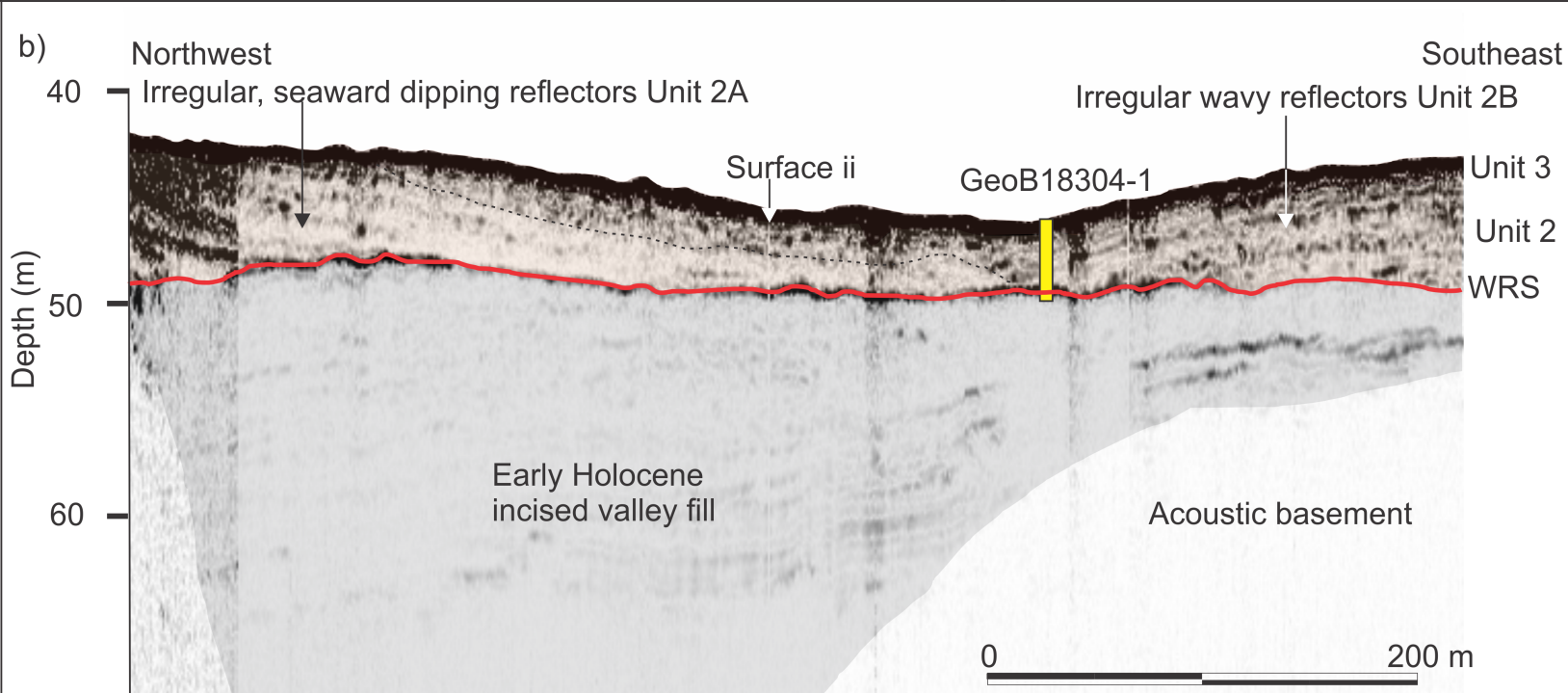
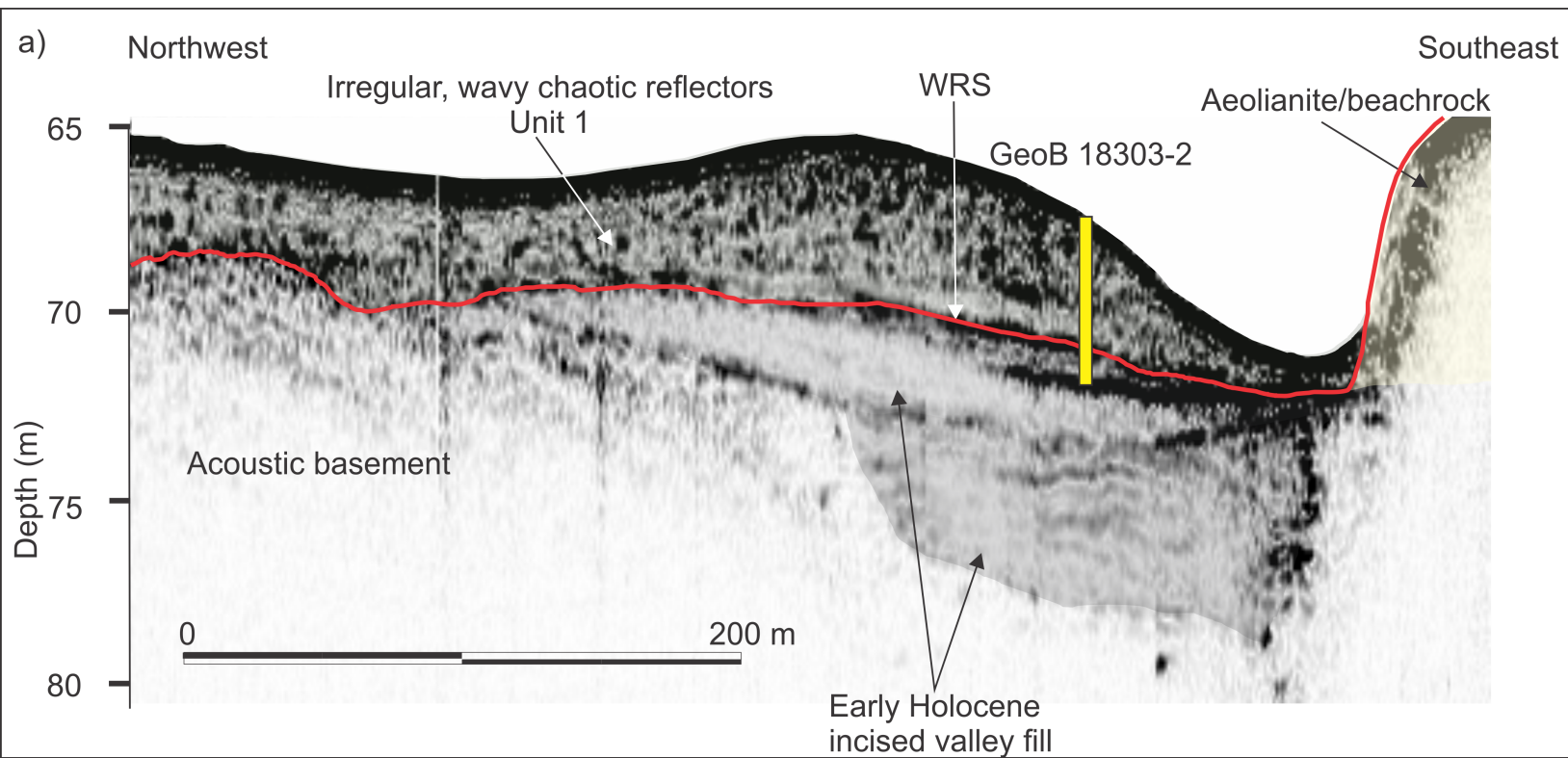
615 **Methods references:**

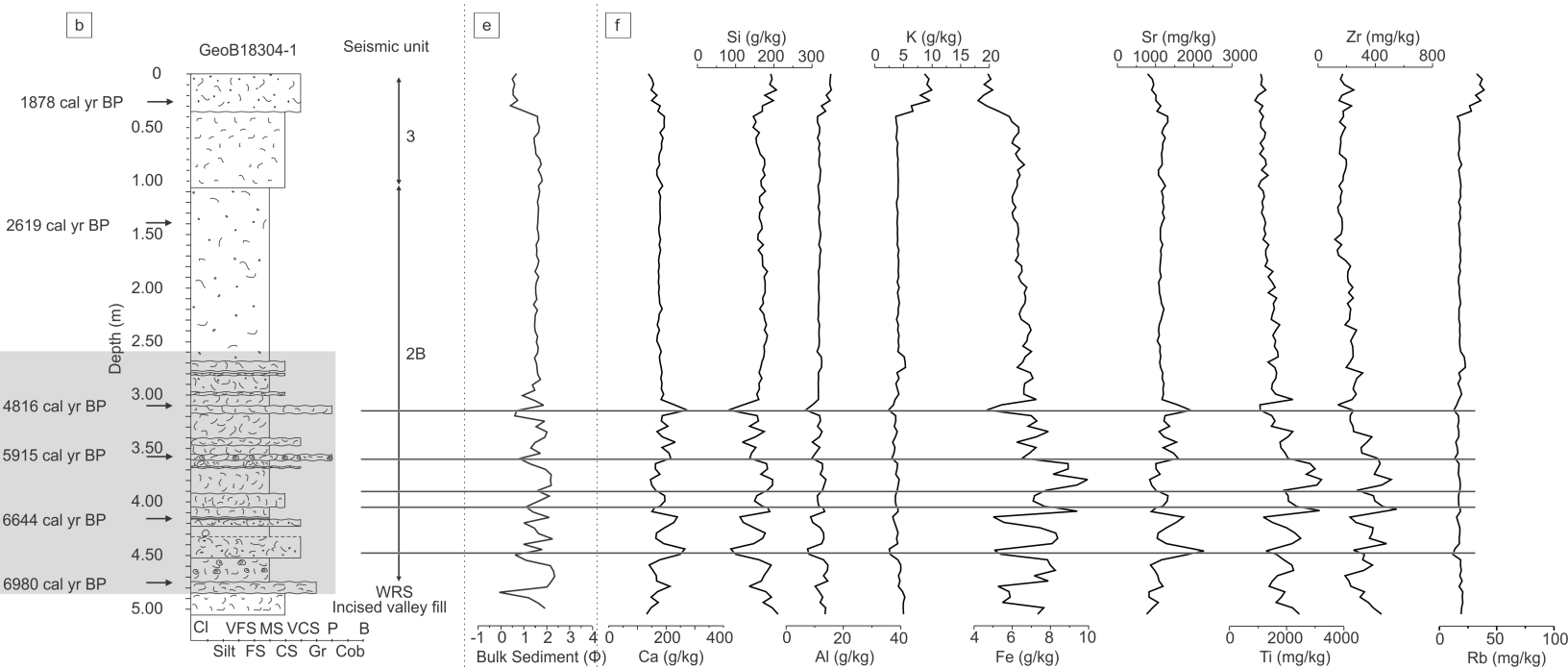
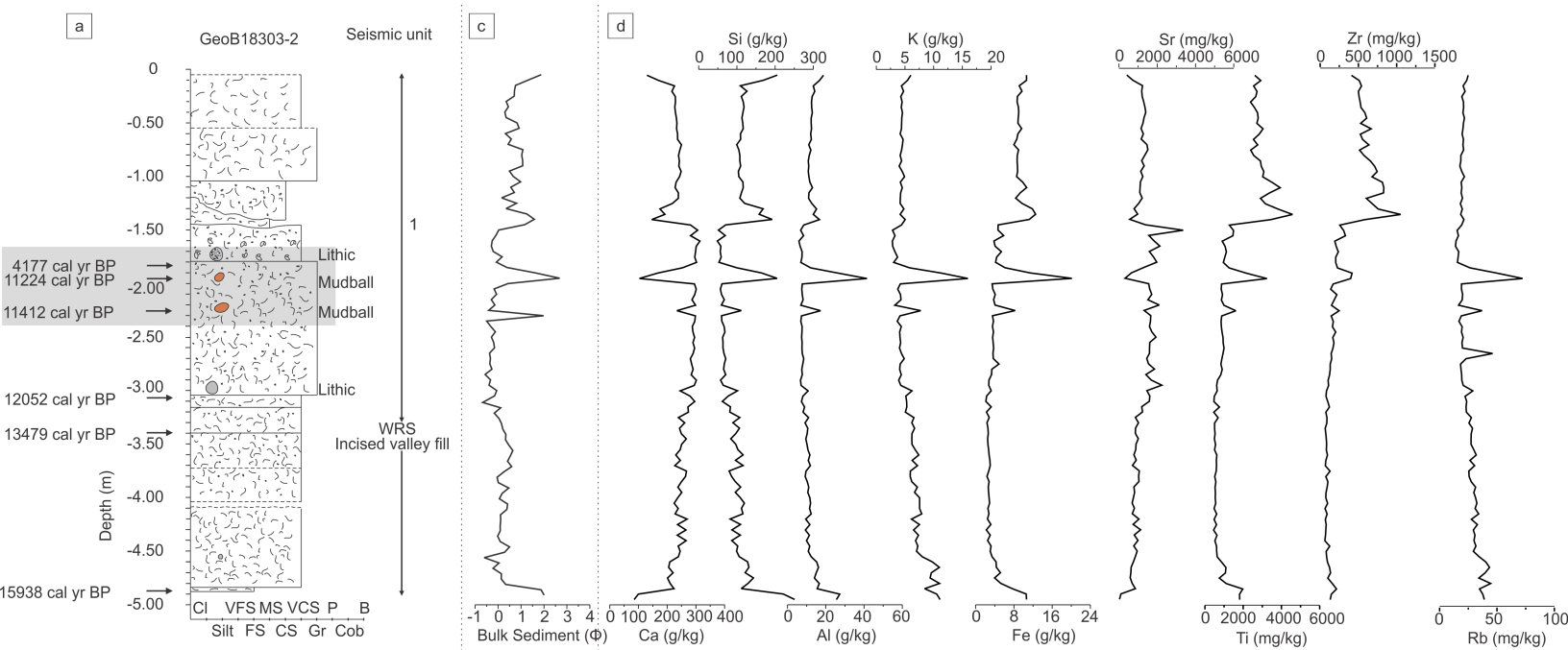
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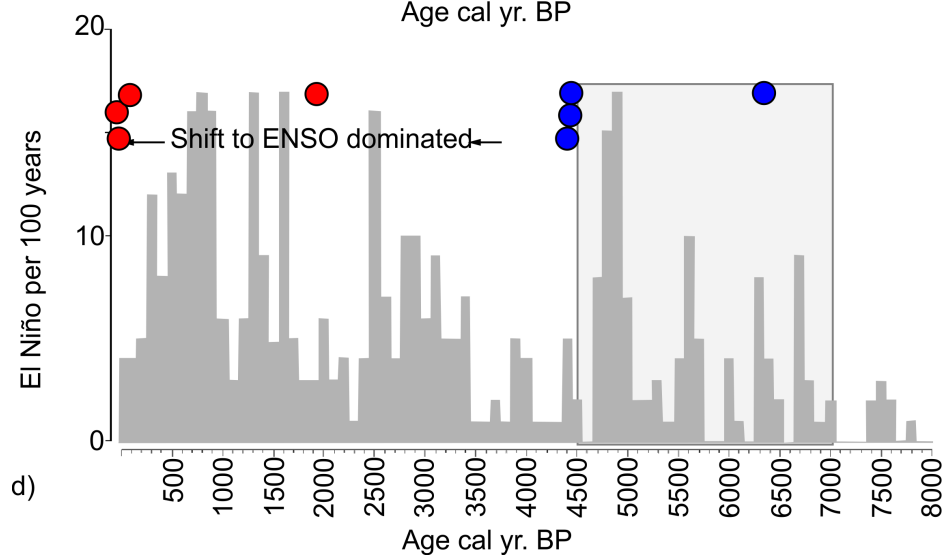
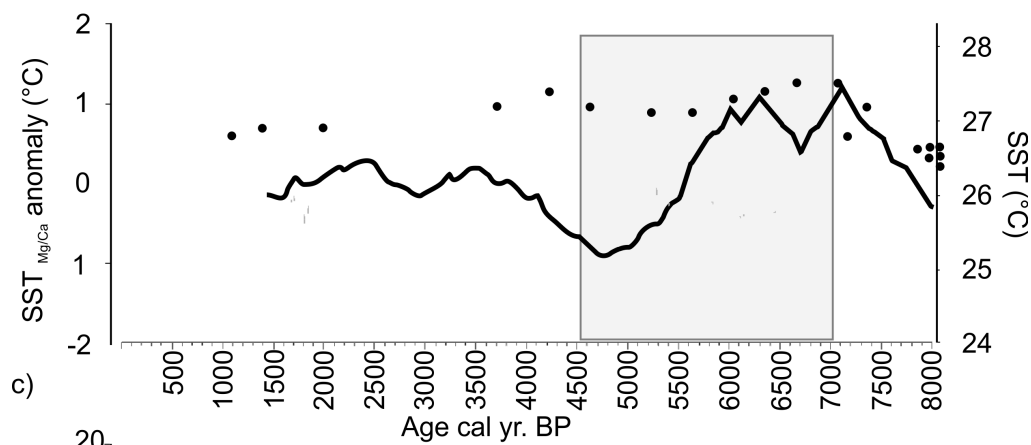
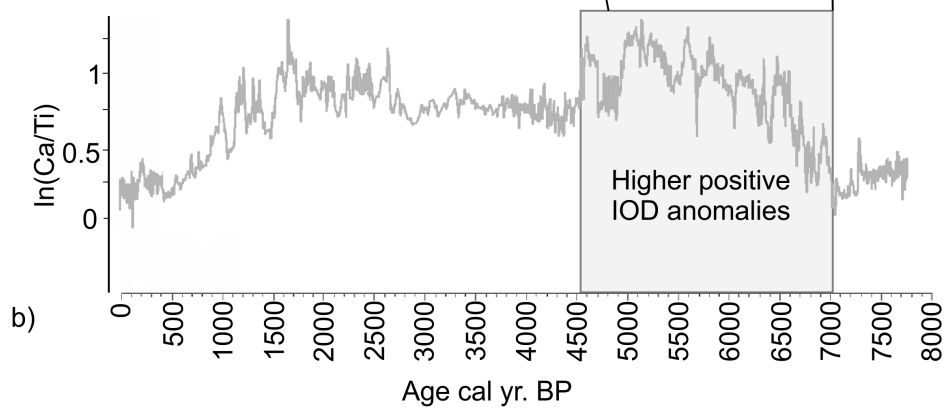
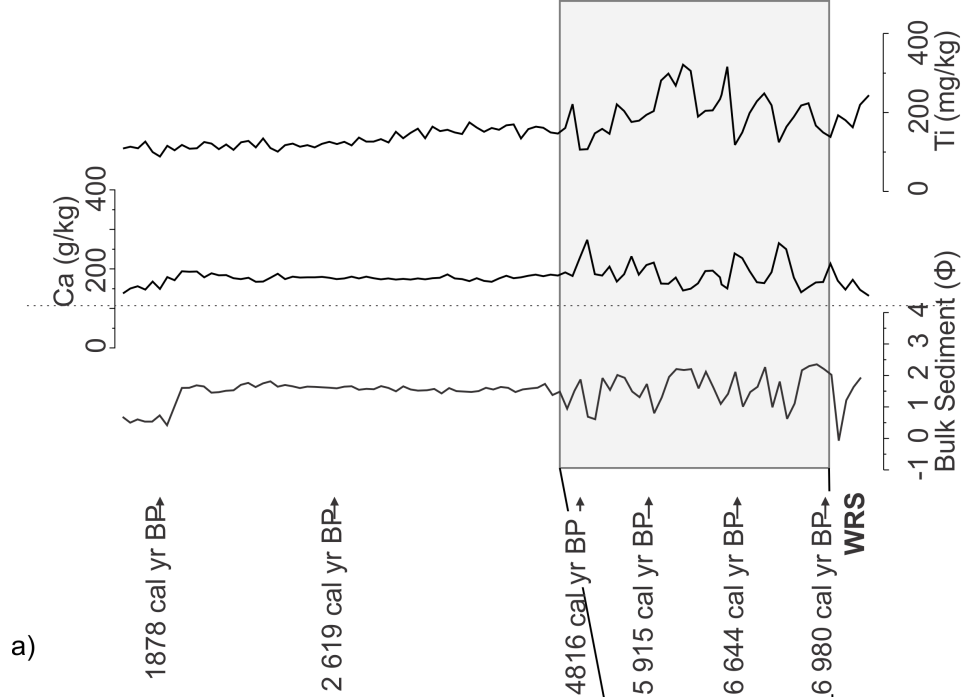
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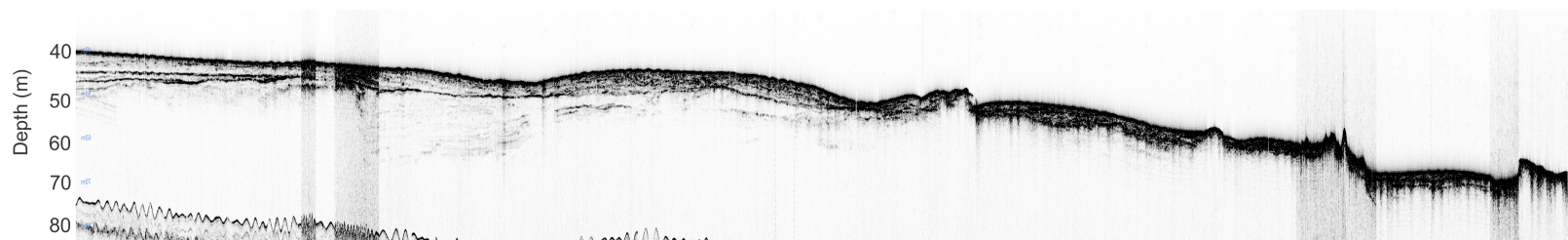
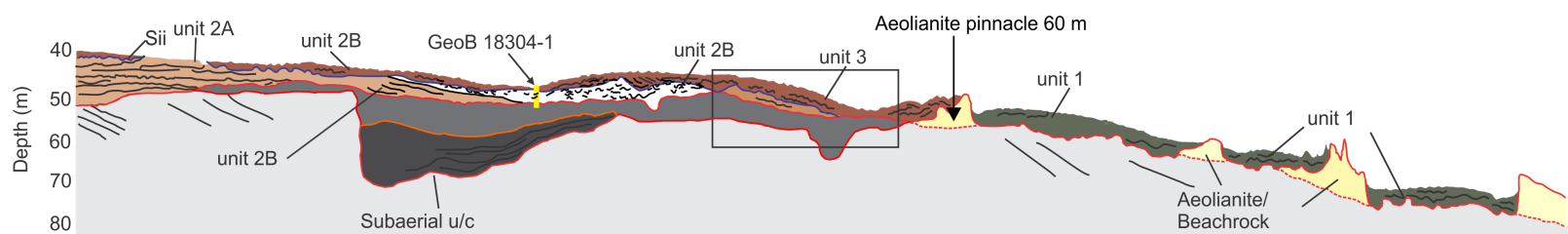
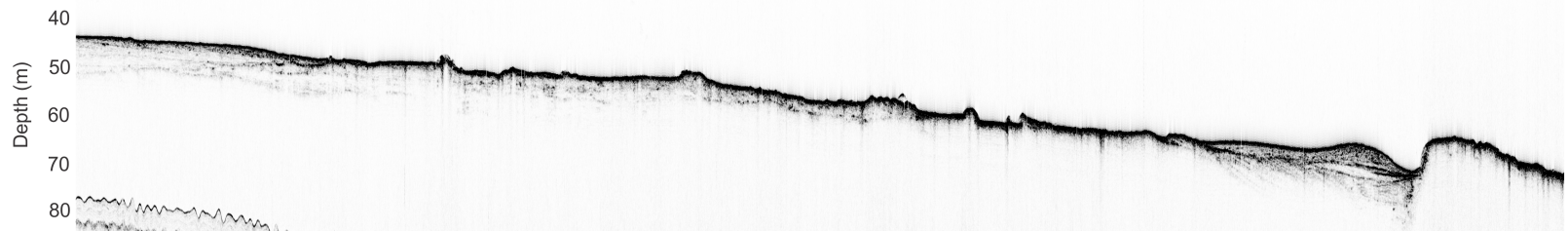
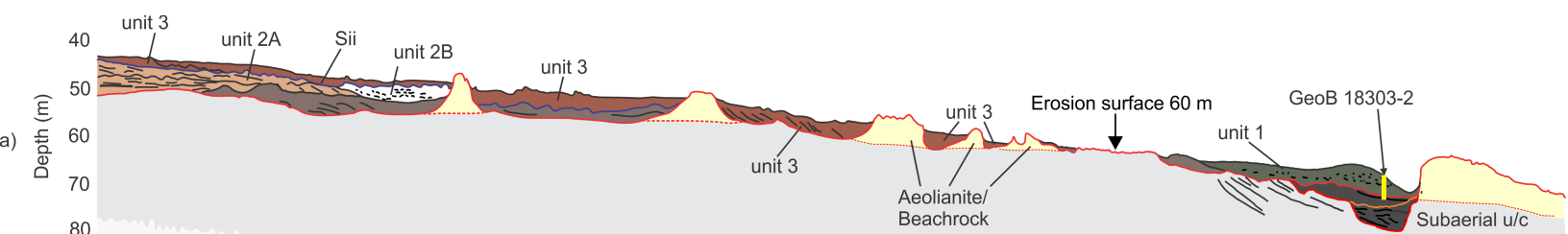
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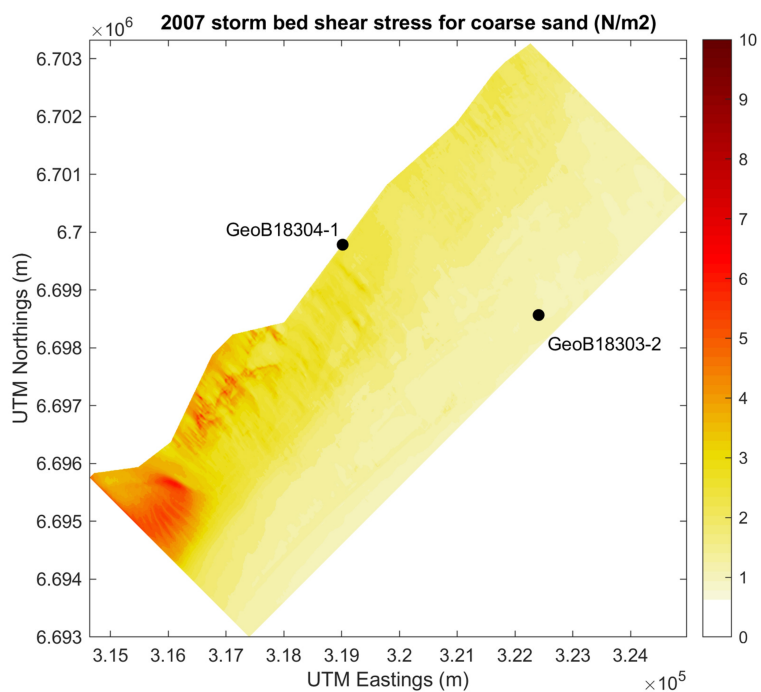




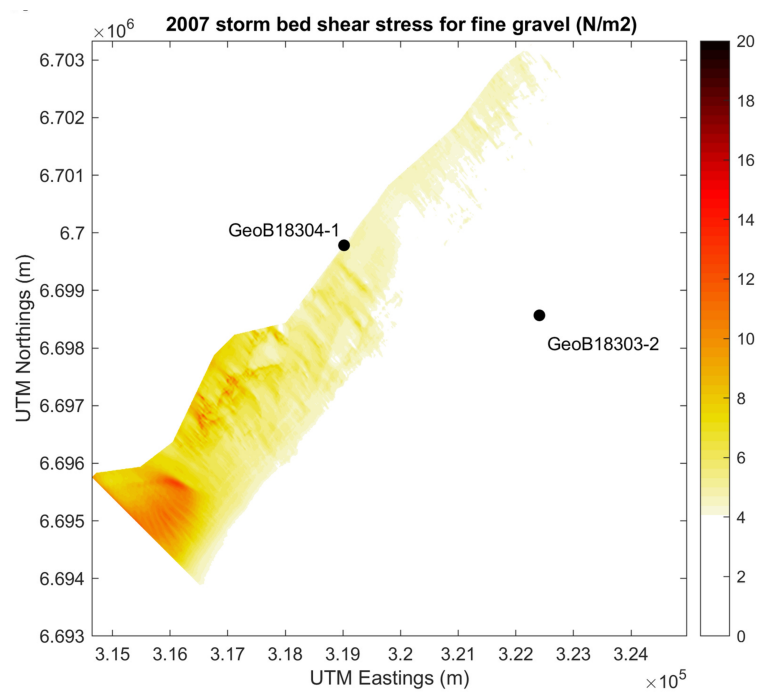




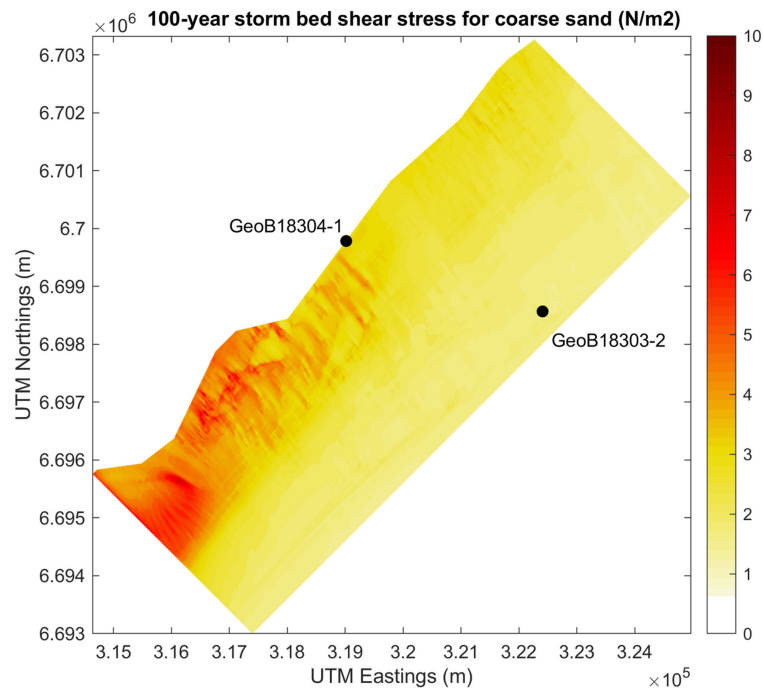
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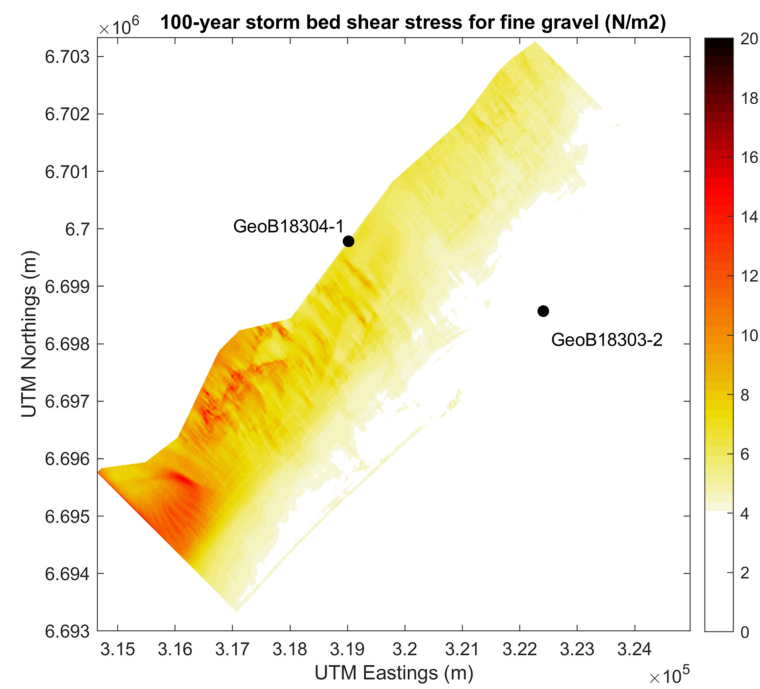
b)



c)



d)



Supplementary Table 2. Chronostratigraphy of GeoB18304-1 and GeoB18303-2. AMS radiocarbon dates are

Depth (cm)	¹⁴ C age yr BP	error ±	Material
Core GeoB18304-1			
25	2 270	30	Bivalve
145	2 845	30	Single gastropod, Nassarius sp
310	4 595	35	Whole shell
359	5 530	40	Articulated bivalve, life position, Eumarcia paupercula
418	6 200	35	Articulated bivalve, life position, Eumarcia paupercula
476	6 480	40	Articulated bivalve, life position, Eumarcia paupercula
Core GeoB18303-2			
190	3 835	35	Bulk organic carbon (outer rim)
190	9 850	50	Bulk organic carbon (centre)
225	10 010	50	Bulk organic carbon
303	10 680	50	CaCO ₃
340	11 690	90	Bulk organic carbon
489	13 300	70	Bulk organic carbon

indicated, together with the composition of material dated and interpretation of the intersected unit/bracketing surfa

Interpretation (Unit/Surface)	Cal age yr BP	
	median	+2σ
Contemporary shoreface, unit 3	1878	1973
Storm-generated gravel/sand couplets, lower shoreface , unit 2	2619	2710
Storm-influenced sand lower shoreface, unit 2B	4816	4910
Storm-generated gravel/sand couplets, lower shoreface, unit 2B	5915	6017
Storm-generated gravel/sand couplets, lower shoreface, unit 2B	6644	6741
Storm-generated gravel/sand couplets, lower shoreface, unit 2B	6980	7127
Exterior of mudball, lower shoreface deposit, unit 1	4177	4383
Interior of mudball, lower shoreface deposit, unit 1	11224	1326
Mudball, lower shoreface deposit, unit 1	11412	11699
Reworked lower shoreface material, overlying wave ravinement surface, unit 1	12052	12346
Incised valley fill, flood tide deltaic package, underlying wave ravinement surface	13479	13583
Incised valley fill, flood tide deltaic package	15938	16180

ice

	Calibration curve
-2σ	
1796	marine 13 (Reimer et al 2013)
2488	marine 13 (Reimer et al 2013)
4695	marine 13 (Reimer et al 2013)
5826	marine 13 (Reimer et al 2013)
6535	marine 13 (Reimer et al 2013)
6870	marine 13 (Reimer et al 2013)
3994	SHcal13 atmospheric curve (Hogg et al. 2013)
11138	SHcal13 atmospheric curve (Hogg et al. 2013)
11244	SHcal13 atmospheric curve (Hogg et al. 2013)
11827	SHcal13 atmospheric curve (Hogg et al. 2013)
13357	SHcal13 atmospheric curve (Hogg et al. 2013)
15707	SHcal13 atmospheric curve (Hogg et al. 2013)